



Frequency-Tunable Pre-stabilized lasers for LISA via Sideband Locking

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7th International LISA Symposium

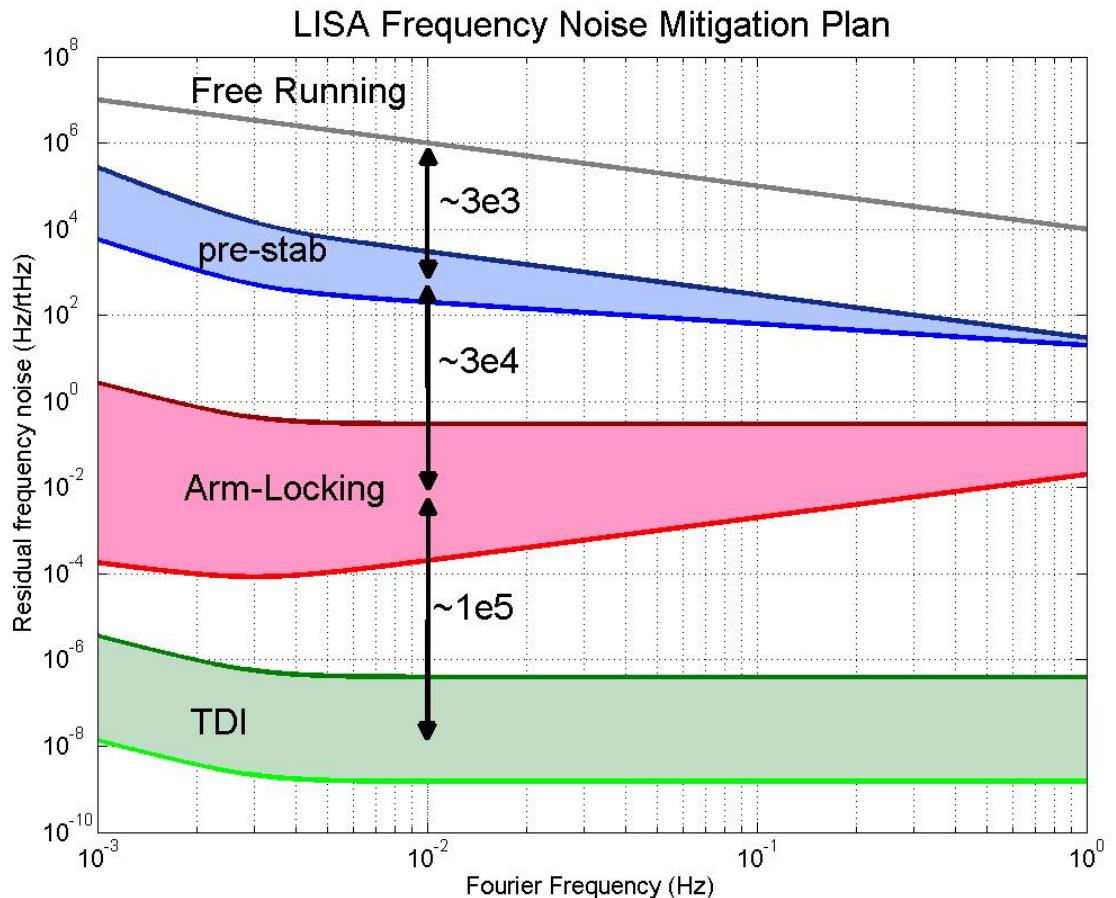
June 18th, 2008

Barcelona, Spain

Beyond Einstein: From the Big Bang to Black Holes

Laser frequency noise is a major potential noise source for LISA

- Three-stage system (two active one passive) to achieve overall suppression of $\sim 10^{13}$
- Running pre-stabilization and arm-locking in series reduces gain (bandwidth) requirements on arm-locking.
- Serial arrangement *requires frequency-tunable pre-stabilization* because cavity lock point drifts with respect to arm-locking lock point.



• Tunable Cavity

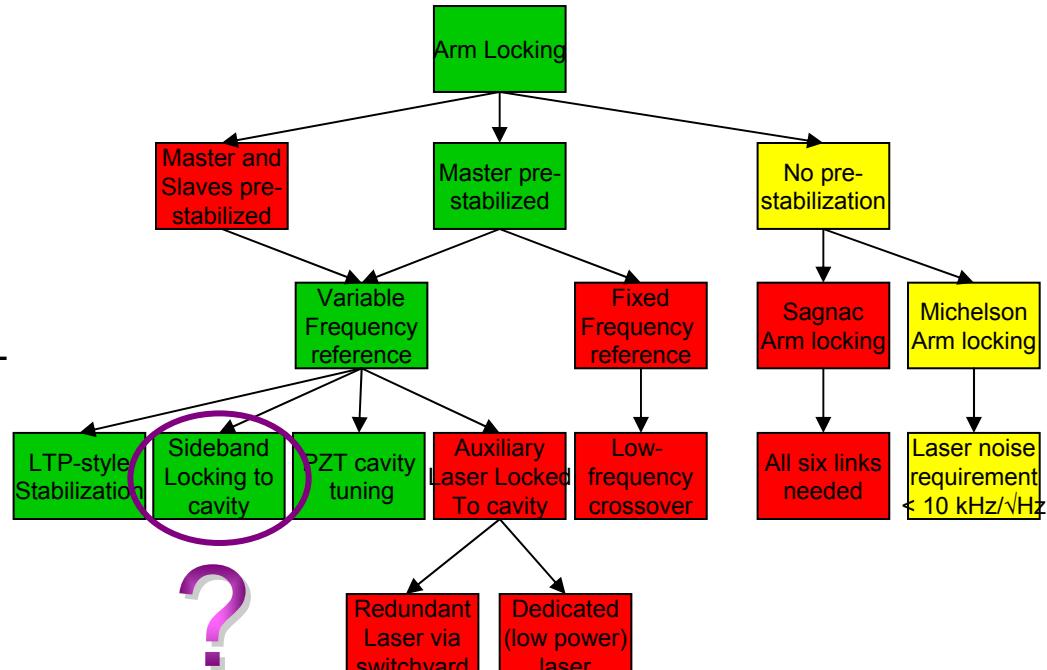
- Adjust length of cavity spacer
- “Standard” method
- Likely reduces cavity stability

• Offset Phase-Lock Loop

- Lock master laser to cavity, PLL slave with adjustable offset
- No modification to cavity
- Requires additional laser

• Offset Sideband Locking

- Use EOM as frequency actuator
- EOM already present for PDH
- No modification to cavity
- Increased noise?

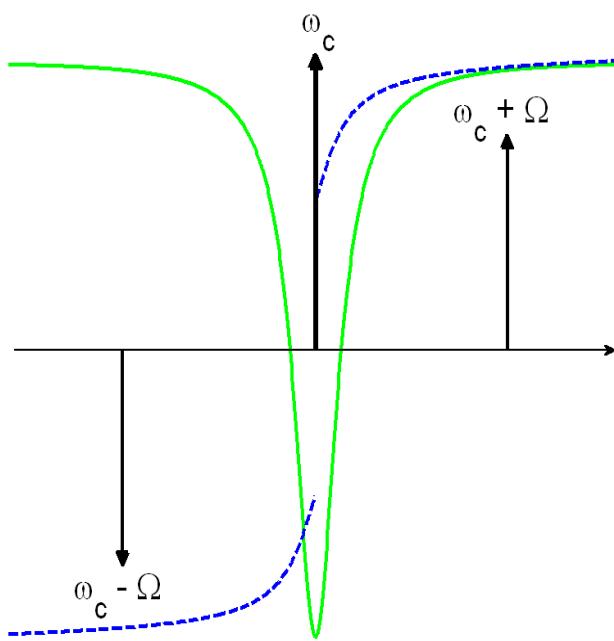


Offset Sideband Locking

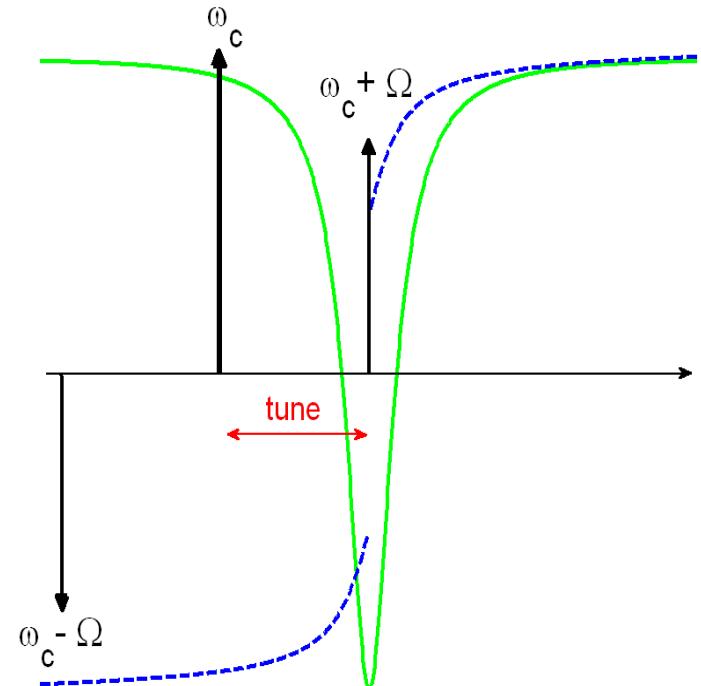
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Concept: Lock phase-modulation sidebands to cavity resonance and tune central frequency by adjusting modulation frequency.

Normal Pound-Drever-Hall Lock

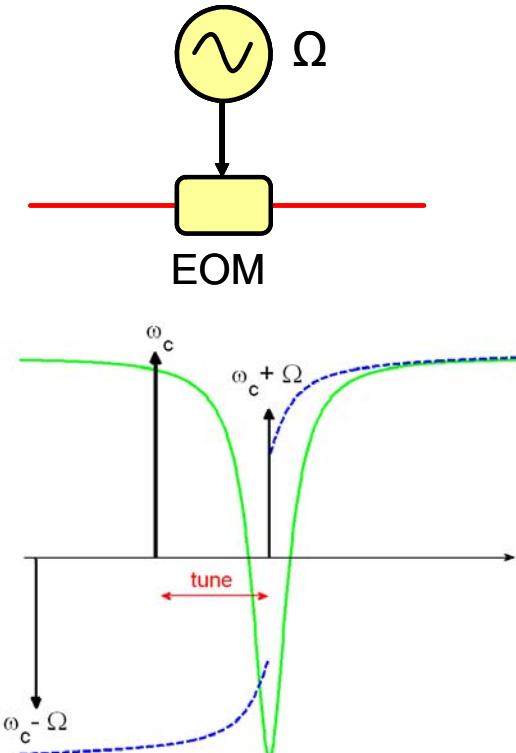


Sideband Lock



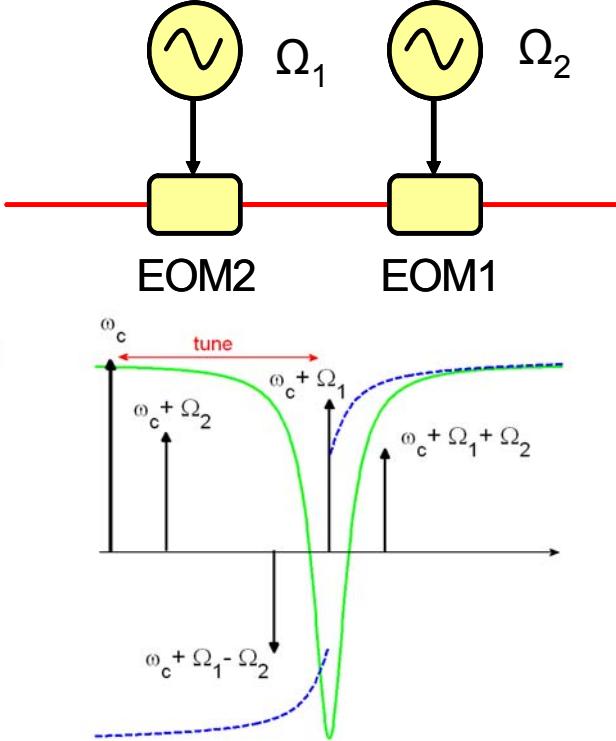
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Single Sideband (SSB)



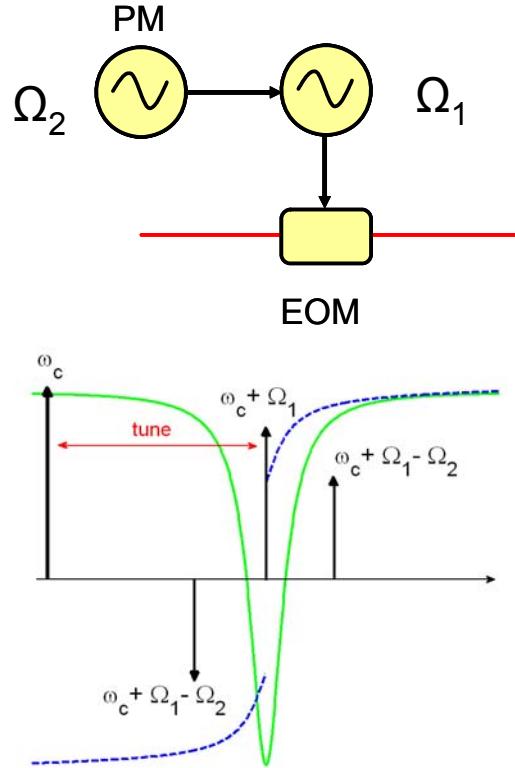
- Simplest to implement
- Some noise coupling due to asymmetry

Dual Sideband (DSB)

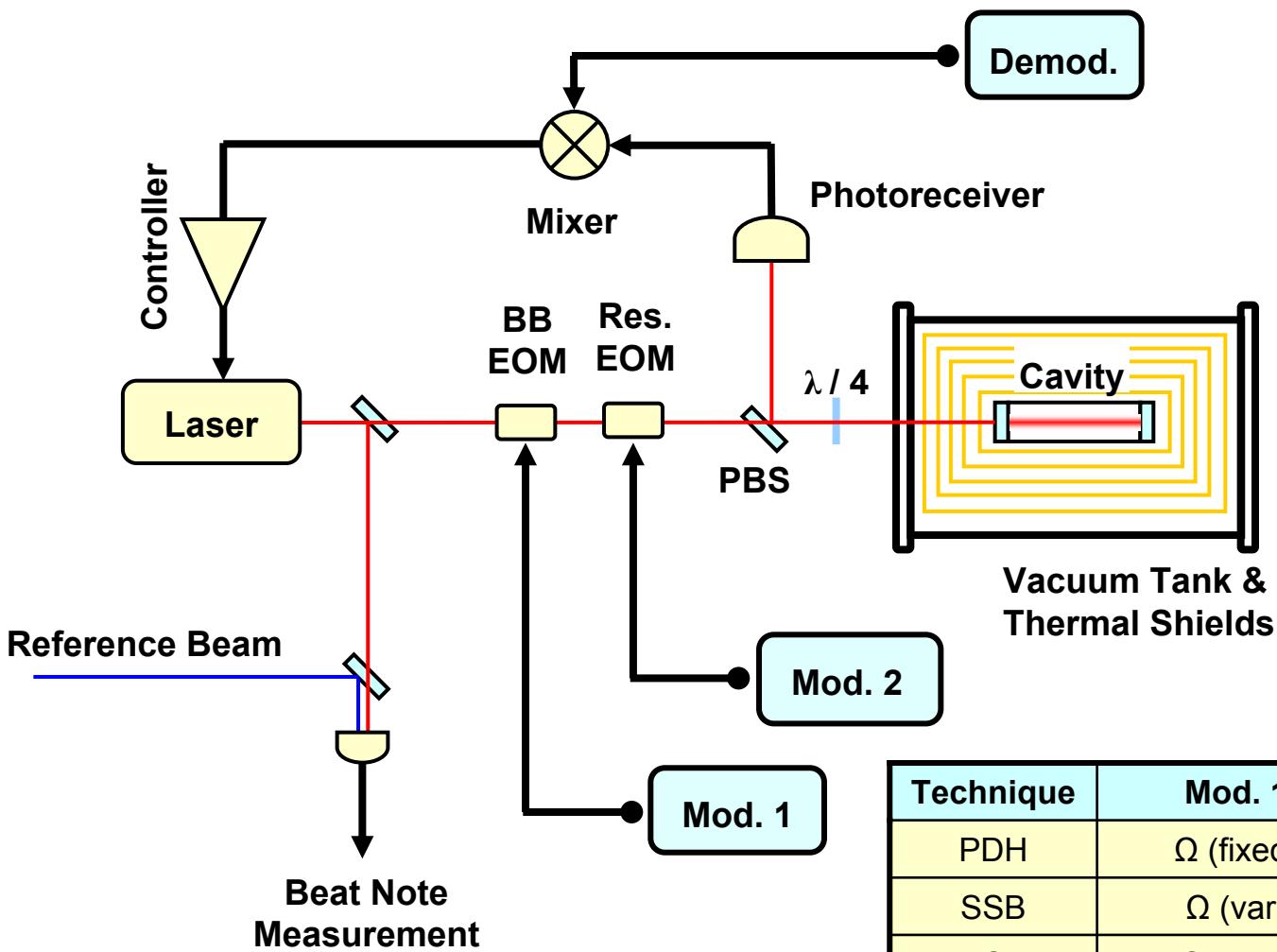


- Restores PDH symmetry
- Complex modulation pattern

Electronic Sideband (ESB)



- Simple, symmetric modulation pattern
- Requires phase modulation capability on LO

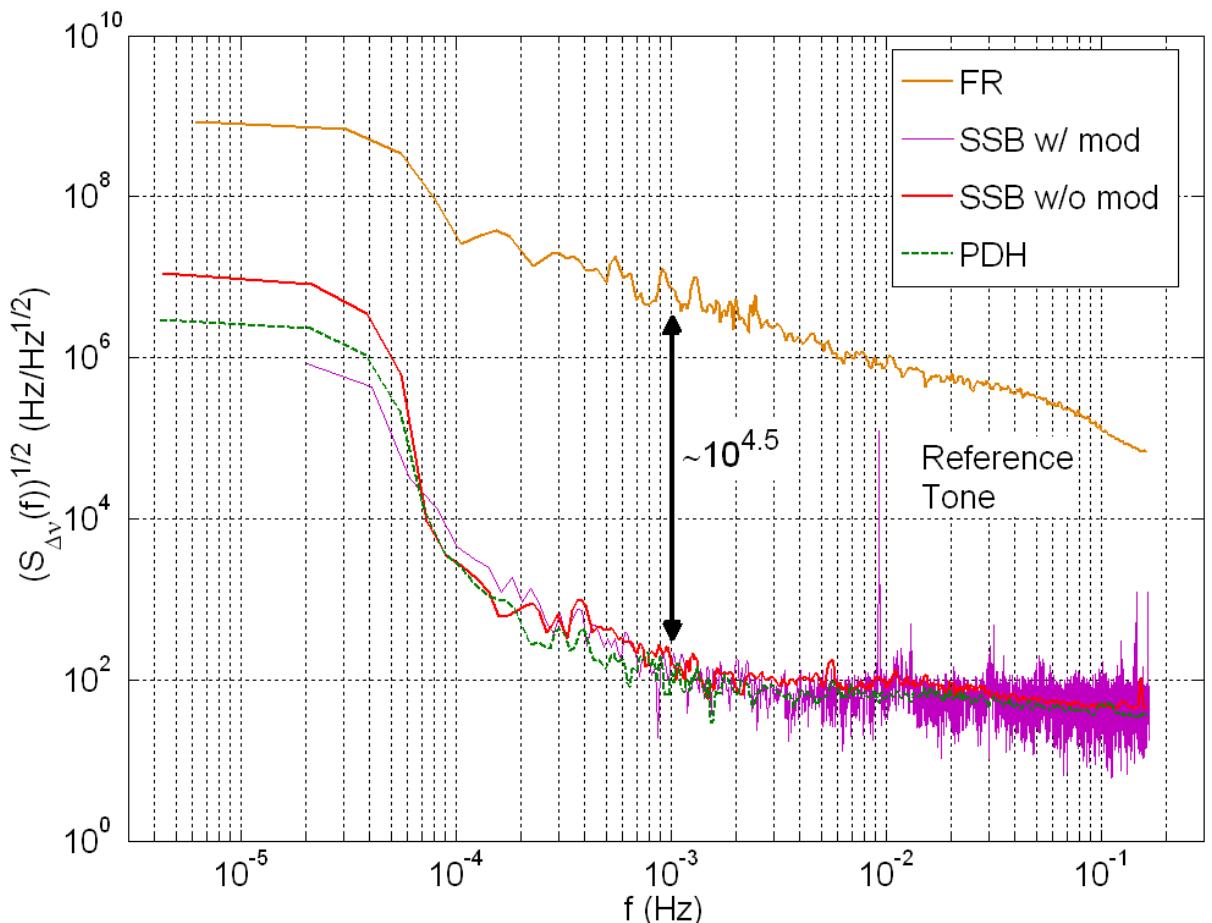


Technique	Mod. 1	Mod. 2	Demod
PDH	Ω (fixed)	N/A	Ω (fixed)
SSB	Ω (var)	N/A	Ω (var)
DSB	Ω_1 (var)	Ω_2 (fixed)	Ω_2 (fixed)
ESB	Ω_1 (var) w/ $\Omega_2 \phi M$ (fixed)	N/A	Ω_2 (fixed)

Stability Results

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- Standard PDH and SSB locking have identical noise performance.
- Adding modulation tone does not disturb the noise floor.
- Similar results with other sideband locking schemes (DSB & ESB)

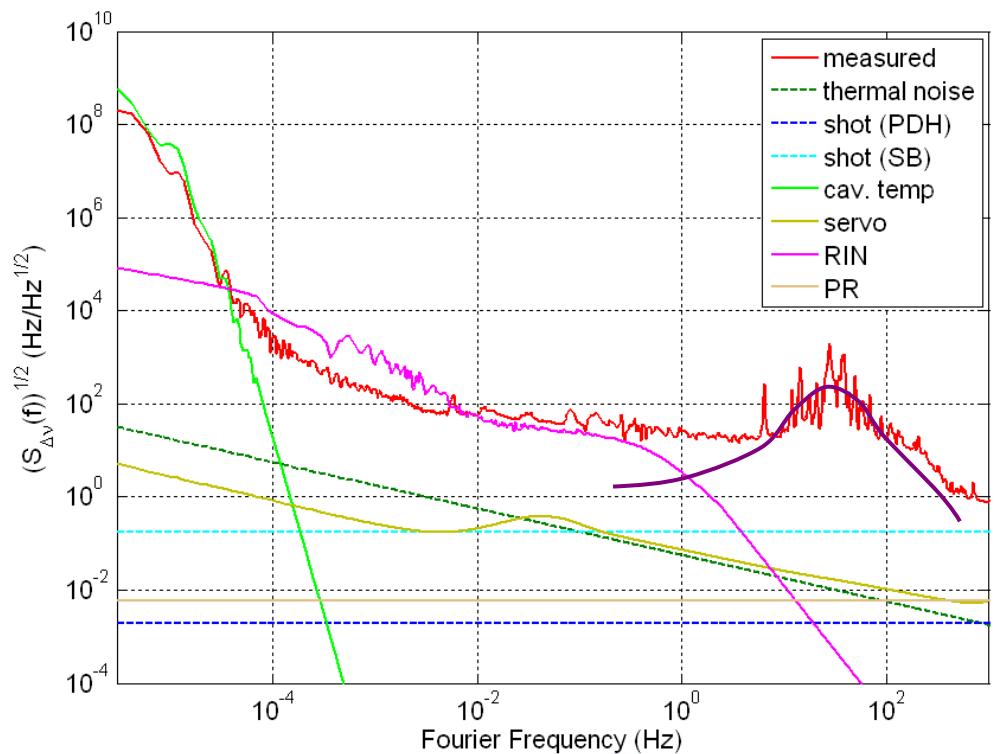


Fundamental Noise

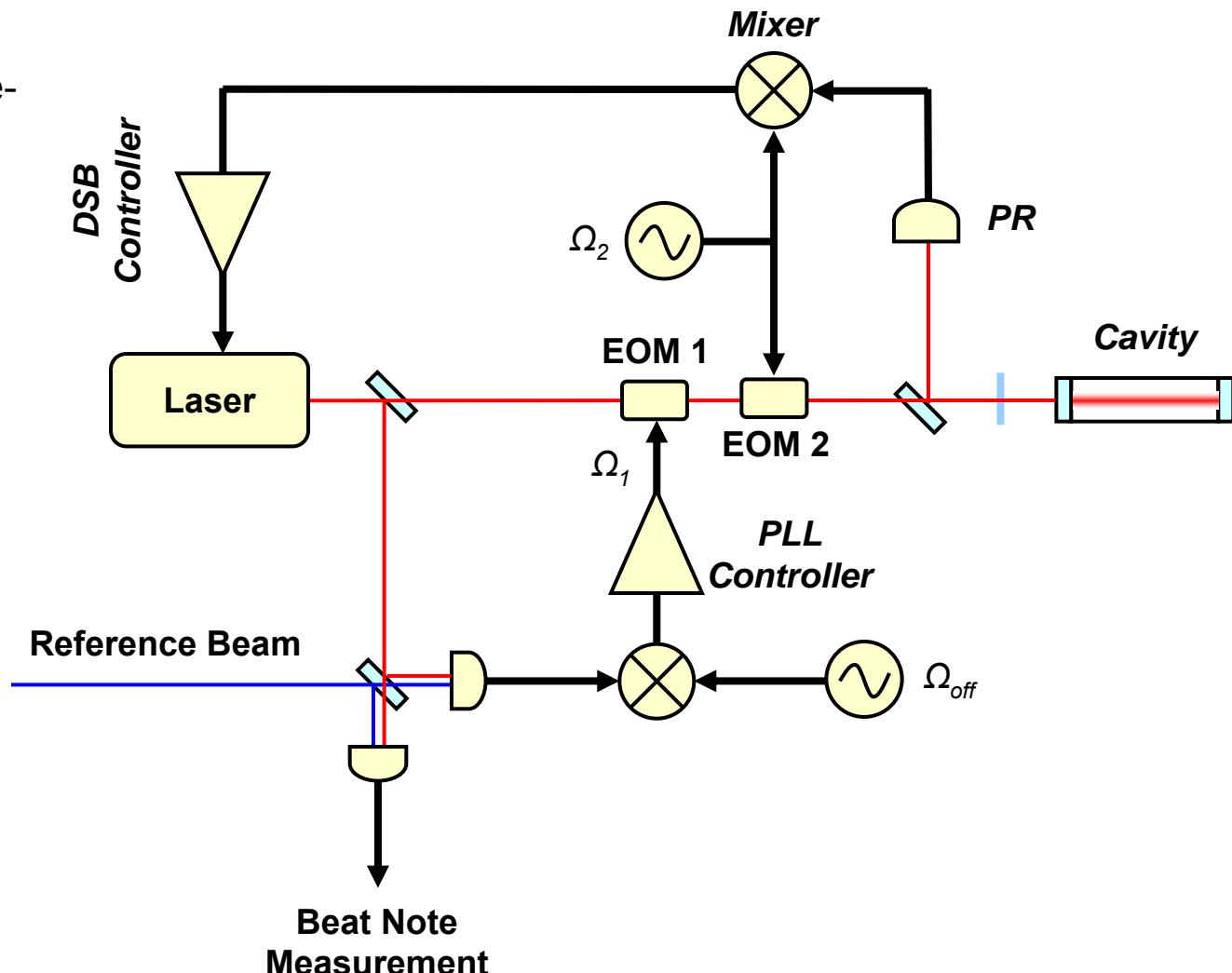
- Shot noise
- Cavity thermal noise

Technical Noise

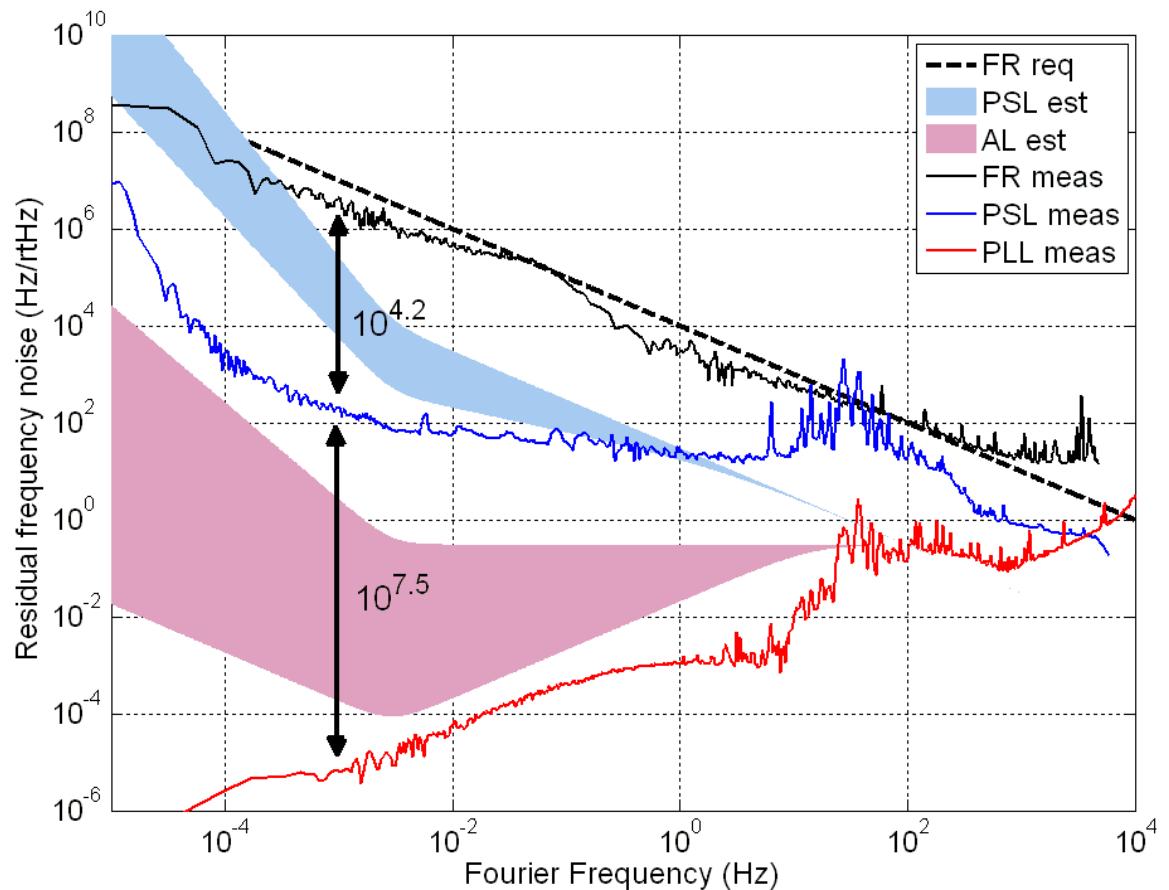
- Temperature Fluctuations
- Servo Noise
- Photoreceiver noise
- RIN
 - via RFAM
 - via absorption
- Vibration Noise/Acoustic
- Pointing
- ???



- Use DSB locking as a pre-stabilization stage in a phase-lock loop
- PLL gain required to achieve given noise floor will be reduced by DSB gain.
- Demonstrates dynamic range, bandwidth, and noise of sideband locking system as an actuator.

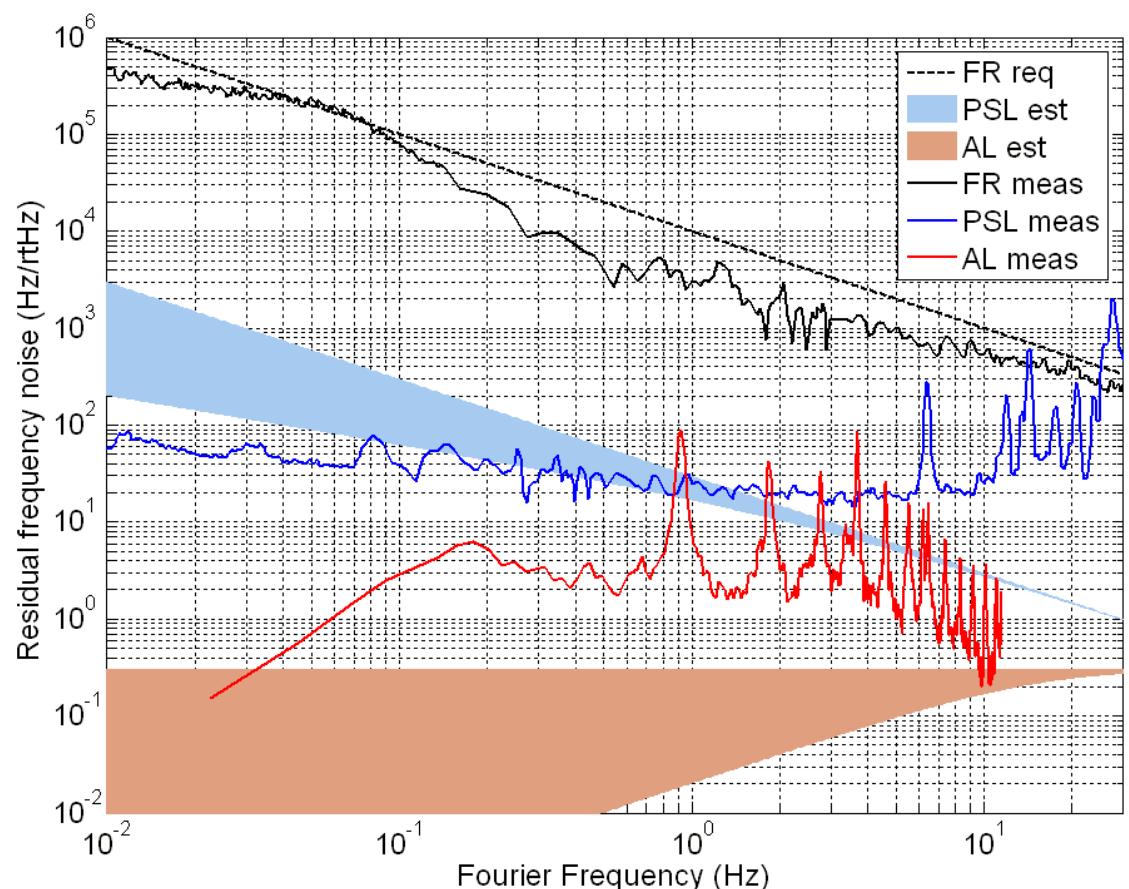


- PSL easily meets requirements below 1Hz
- PSL noise high above 1Hz, believed to be vibration/acoustics & RIN
- PLL noise floor beats arm-locking CBE, indicating actuator noise is sufficiently small for arm-locking



- Simulate 1-s long arm using EPD technique
- Perform single-arm locking in place of PLL
- Observe characteristic arm-locking noise spectrum

See V. Wand's talk
 &
 Y. Yu's poster



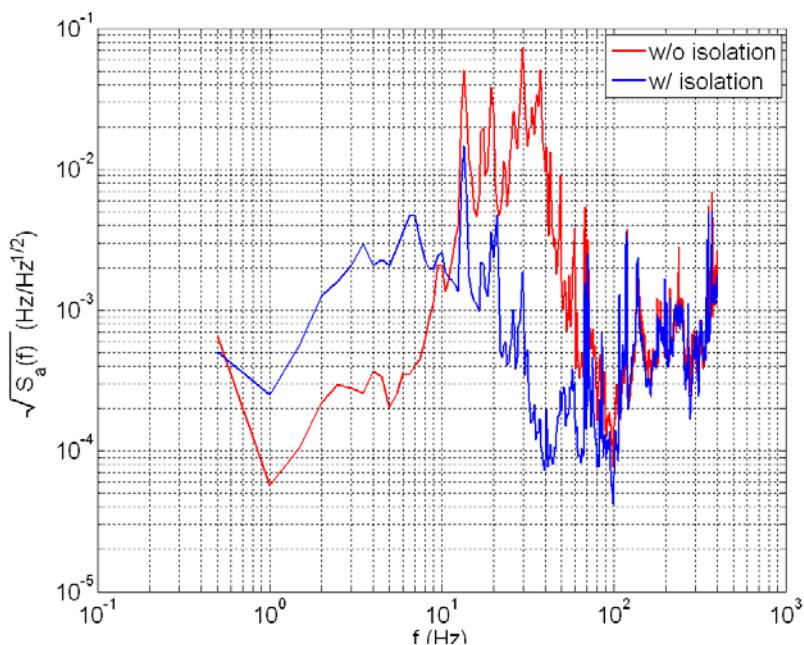
- Current LISA Frequency Noise Plan requires frequency-tunable pre-stabilization system
- Offset Sideband locking can provide frequency tuning with minimal modification to standard cavity-locking technique
- Laboratory results demonstrate that sideband locking can meet the LISA pre-stabilization requirements
- Initial demonstration of first two steps to LISA frequency noise suppression.



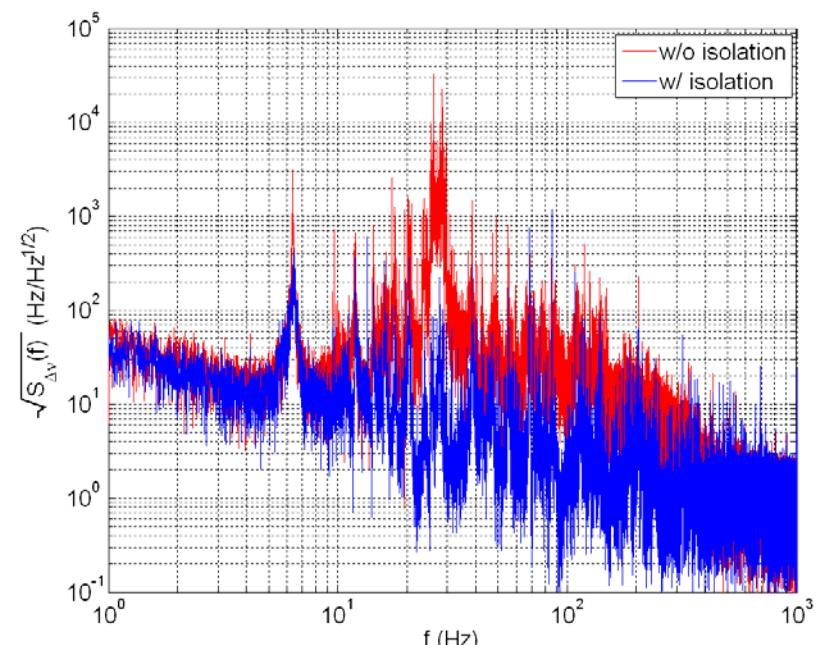
Backup Slides

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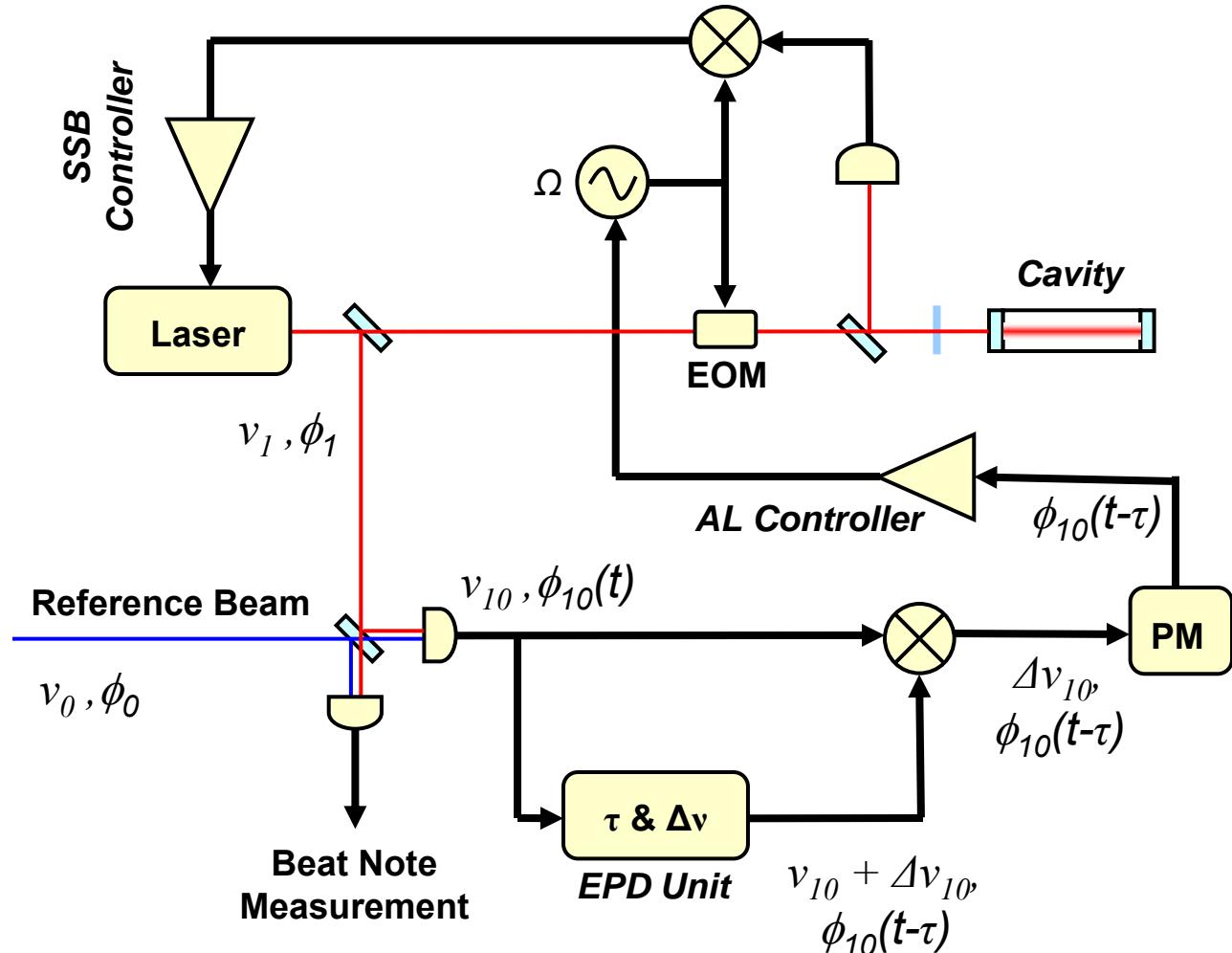
Table Acceleration Noise
Measured via Accelerometer



Frequency Noise measured
via phasemeter

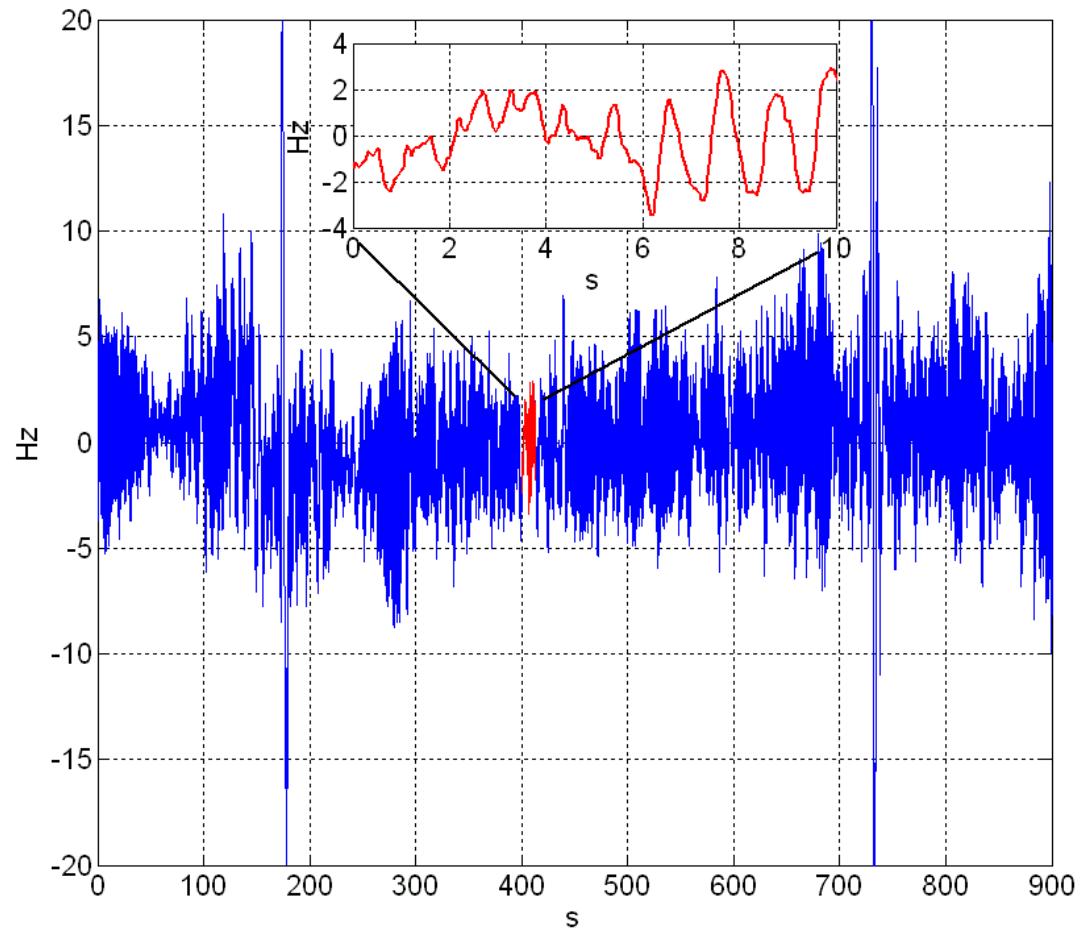


- Extension of PLL example utilizing Electronic Phase Delay (EPD) technique to mimic arm-locking error signal.
- Simulates single-arm locking
- $\sim 1\text{s}$ round-trip delay
- Fixed Doppler shift
- Utilized SSB locking for first test.



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- 1 s delay gives noise at 1 Hz multiples
- Pre-stabilization noise is non-stationary, glitches excite arm-locking transients



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- Measured noise suppression matches expectations
- ~40dB at 100mHz

